

The Brightest Galaxies at Cosmic Dawn



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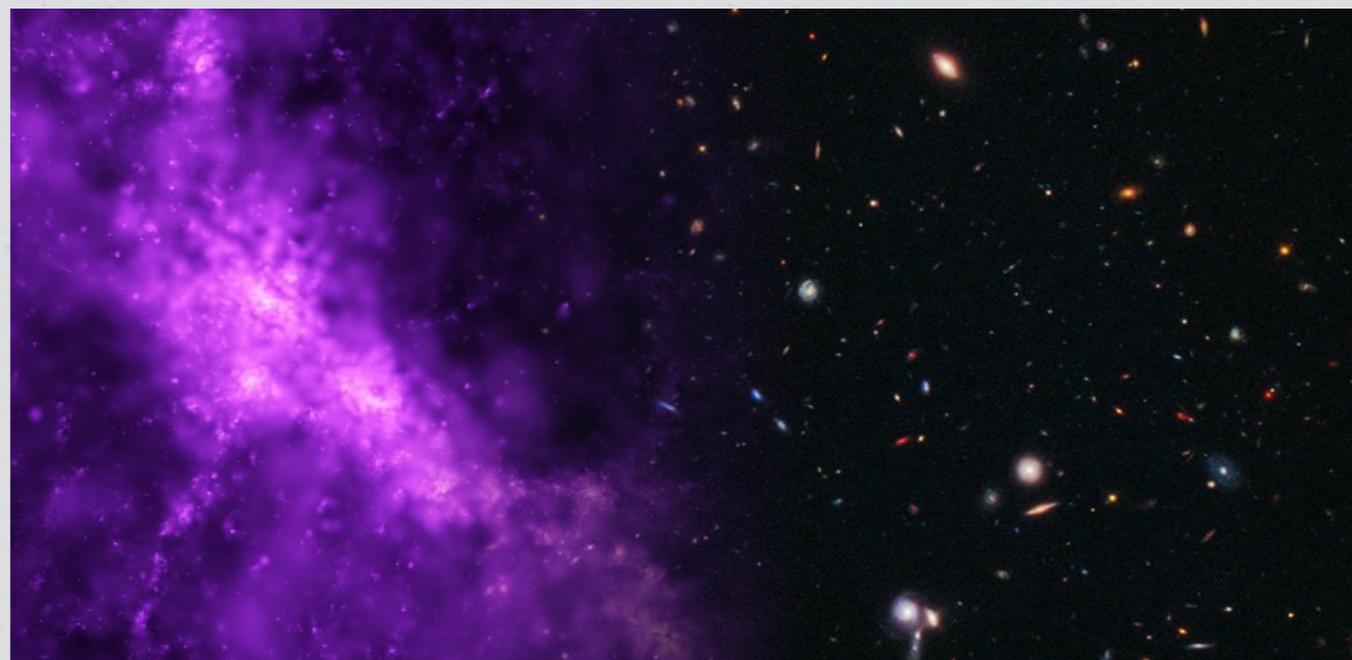
Australian Government
Australian Research Council

with Stephanie Bernard, Charlotte Mason, Daniela Carrasco, Valentina Calvi, Keven Ren

& the



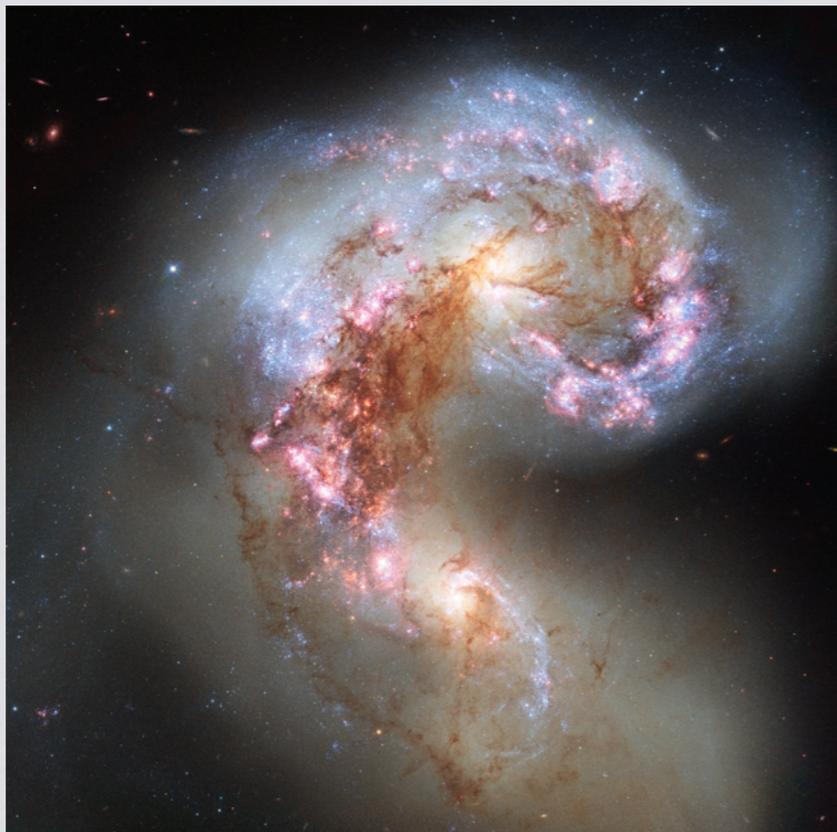
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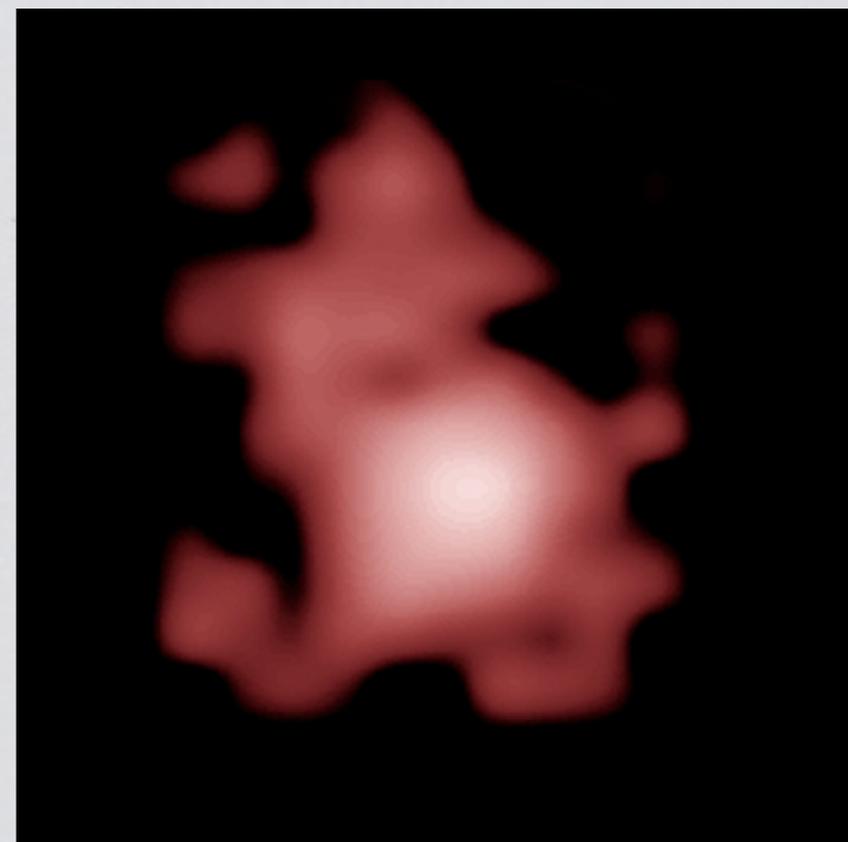
Aspen - March 11th 2016

Star formation: now and then

★ Was star formation different?



today



~13.4 Gyr ago

Image credits: NASA/ESA

Galaxies in the first 700Myr

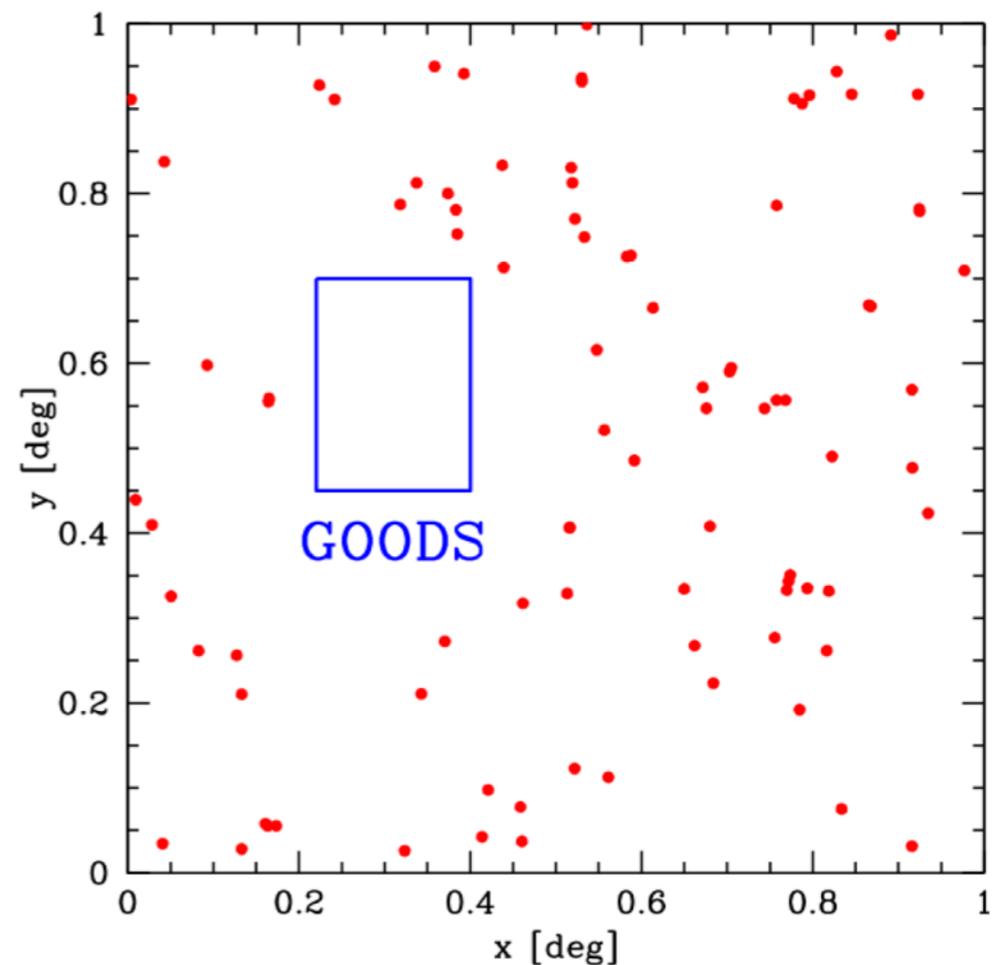
- Exciting results from Hubble legacy fields

★ Legacy fields challenges:

★ (Ultra)Deep, small area:
Mostly faint galaxies ($L < L^*$)

★ Few lines of sight:
Large area surveys
affected by galaxy
clustering

Bright $z \sim 9$ galaxies (AB mag < 27)
Simulated clustering



The Brightest of Reionizing Galaxies Survey (Trenti+ 2011)



Key goal: Identify bright galaxies at $z \sim 8-10$

- ★ Optical+near-IR WFC3 pure parallel imaging:
V + Y,J,[JH],H ($m_{AB} \sim 26.5-27 @ 5\sigma$)
- ★ 900 orbits (~ 60 days) since 2010:
>150 WFC3 **independent** fields, >700 arcmin²
- ★ 21 diverse peer-reviewed publications:
Dataset has legacy value from $z \sim 10$ to $z \sim 0$
- ★ Public data release through MAST

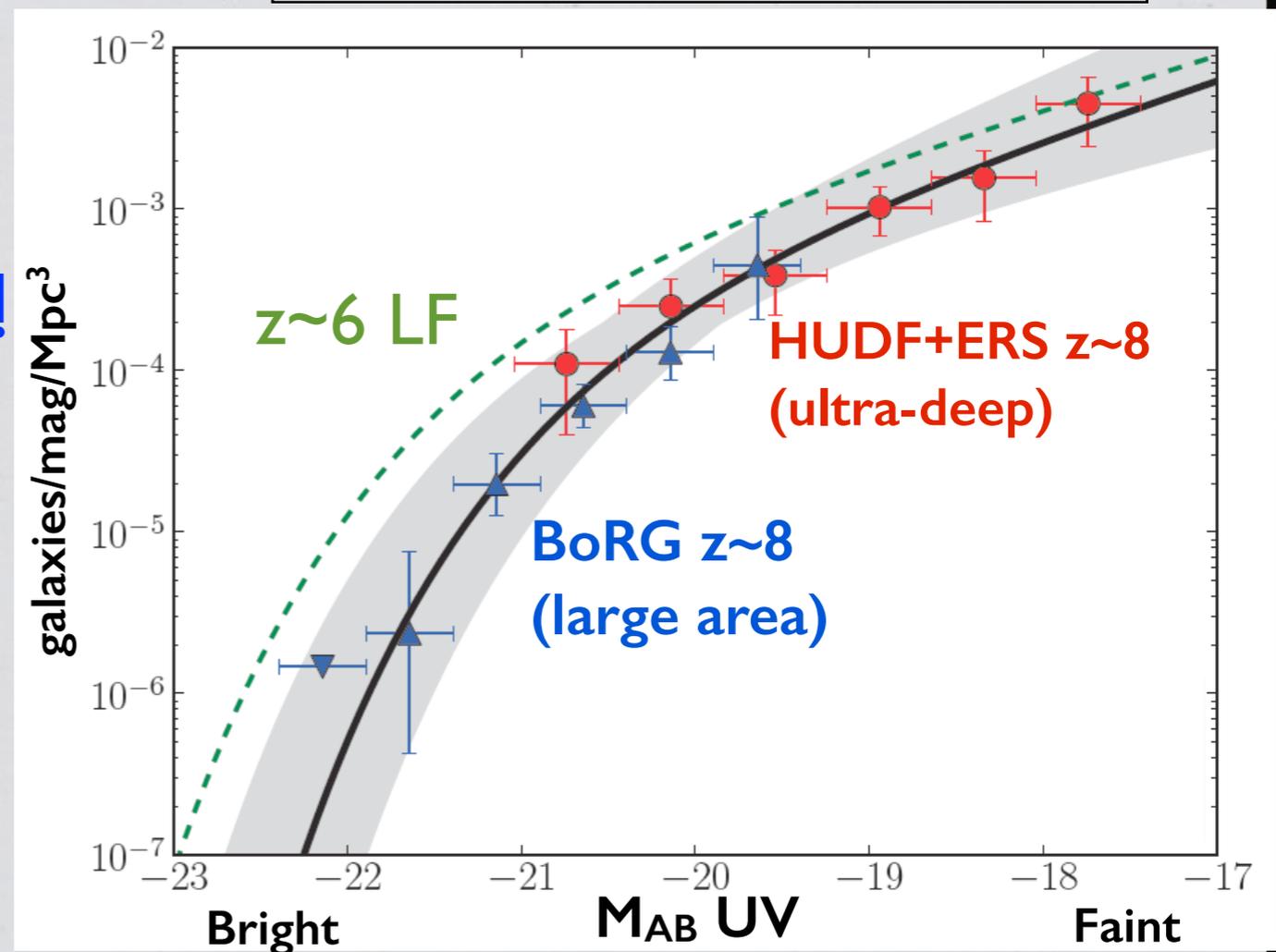
2010-14: The luminosity function

at $z \sim 8$

Large area ($\sim 350 \text{ arcmin}^2$) determination [2014]

- ★ BoRG+HUDF/ERS:
97 Y-dropout galaxies
- ★ None known preWFC3!
- ★ LF well described by Schechter form
- ★ Tension with ground-based $z \sim 7$ results?
(Bowler's talk Monday)

$$\phi(L) = \phi_0(L/L_*)^\alpha \exp(-L/L_*)$$

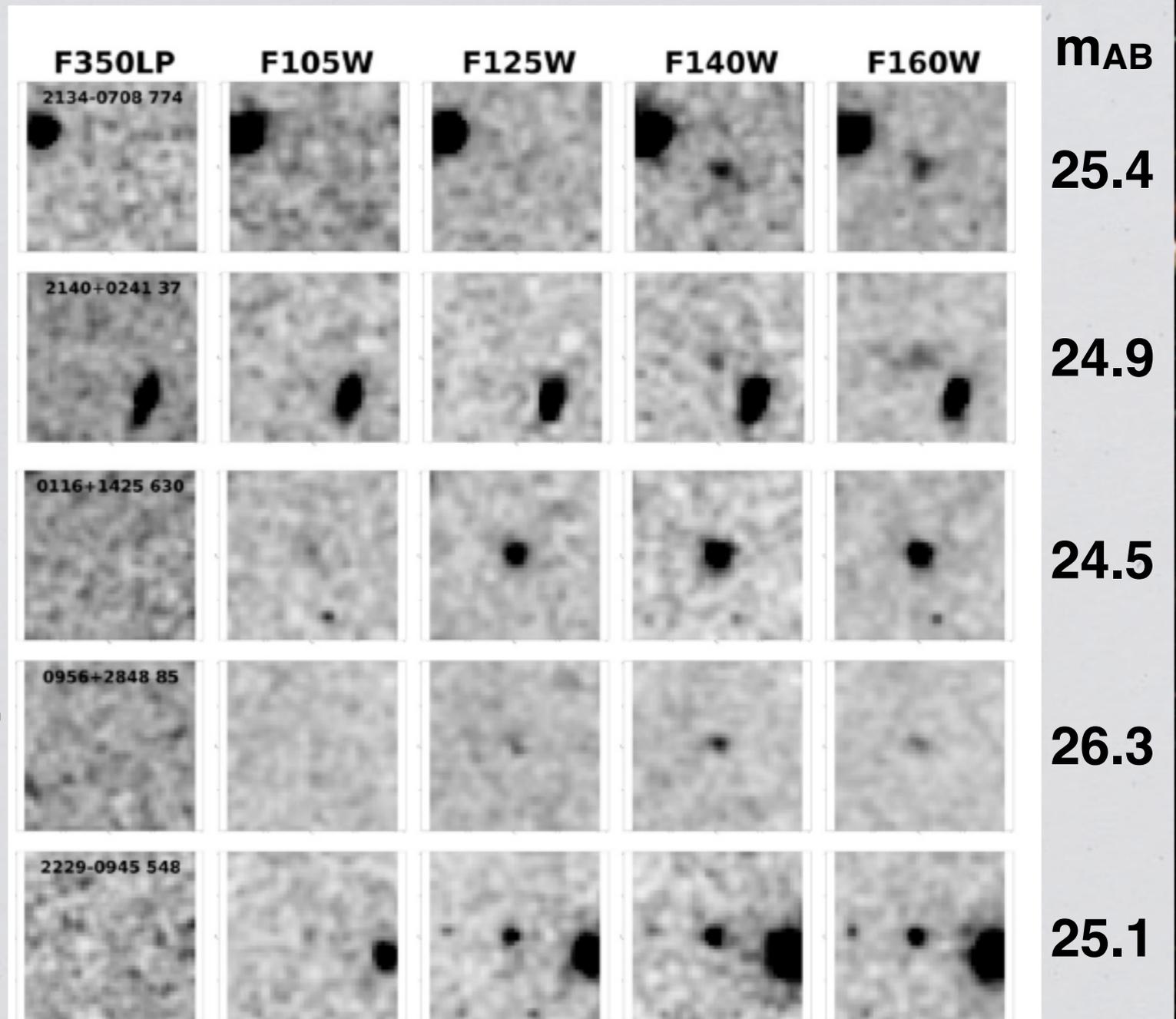


Bradley, MT et al. (2012); Schmidt, MT et al. (2014)



2016: The new frontier: $z \sim 8-11$

- ★ Several $z \sim 8.5-11$ candidates ($m_{AB} \sim 24.5-26$) from 130 arcmin² of new cycle 22 data
- ★ Full cycle 22 dataset analysis ongoing (~ 350 arcmin²)

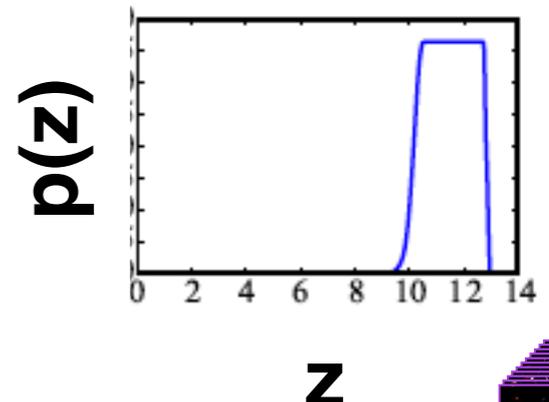
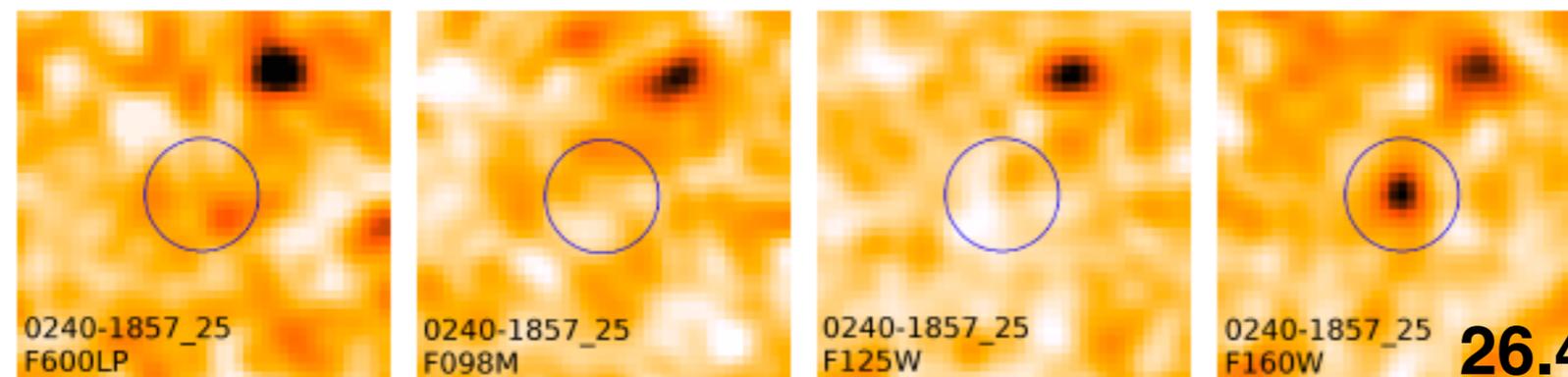
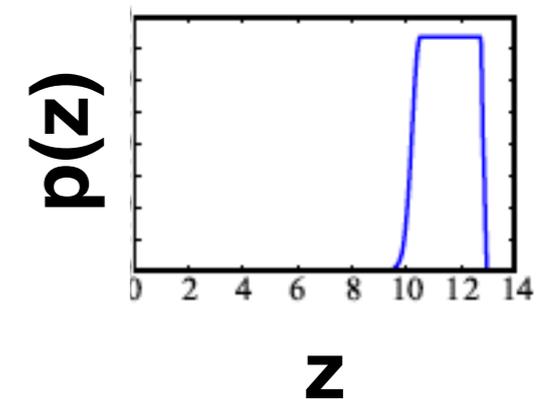
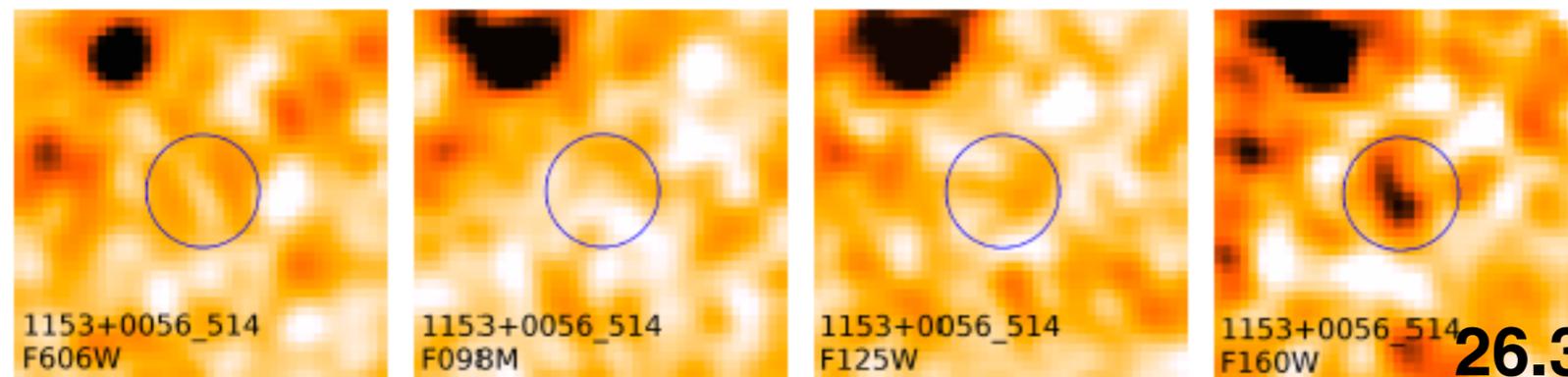
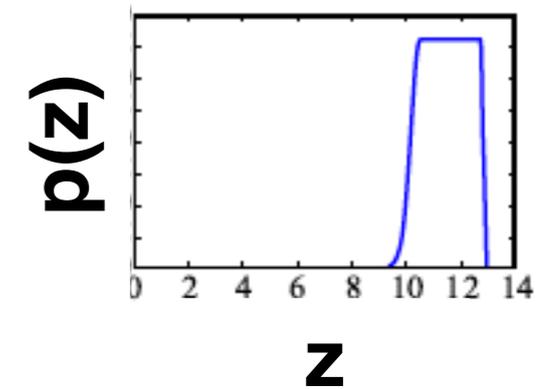
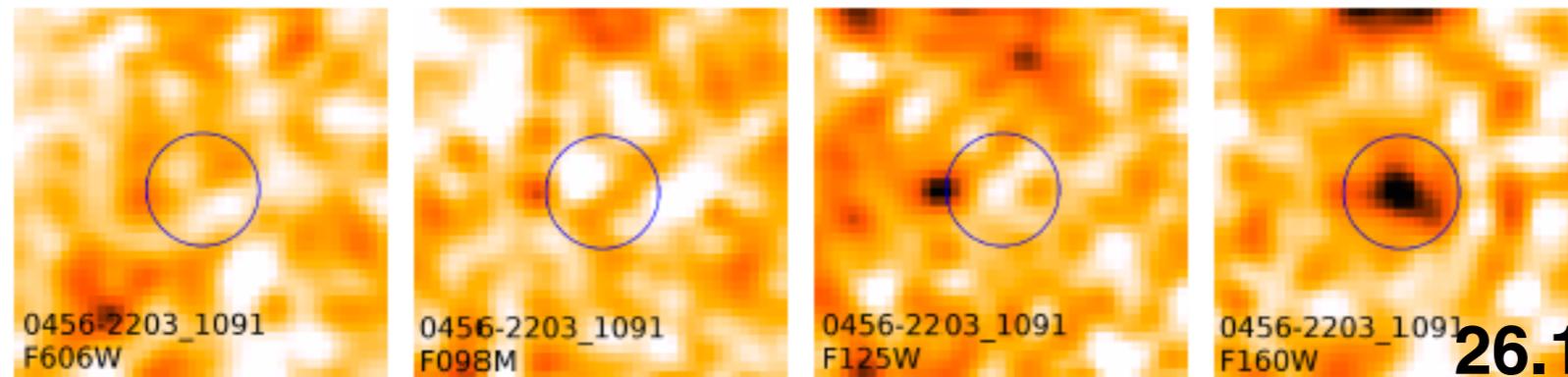


Calvi et al. (2016)



2016: The new frontier: $z \sim 10$

- Analysis of archival BoRG yields $z \sim 10$ catches



Bernard et al. (submitted)

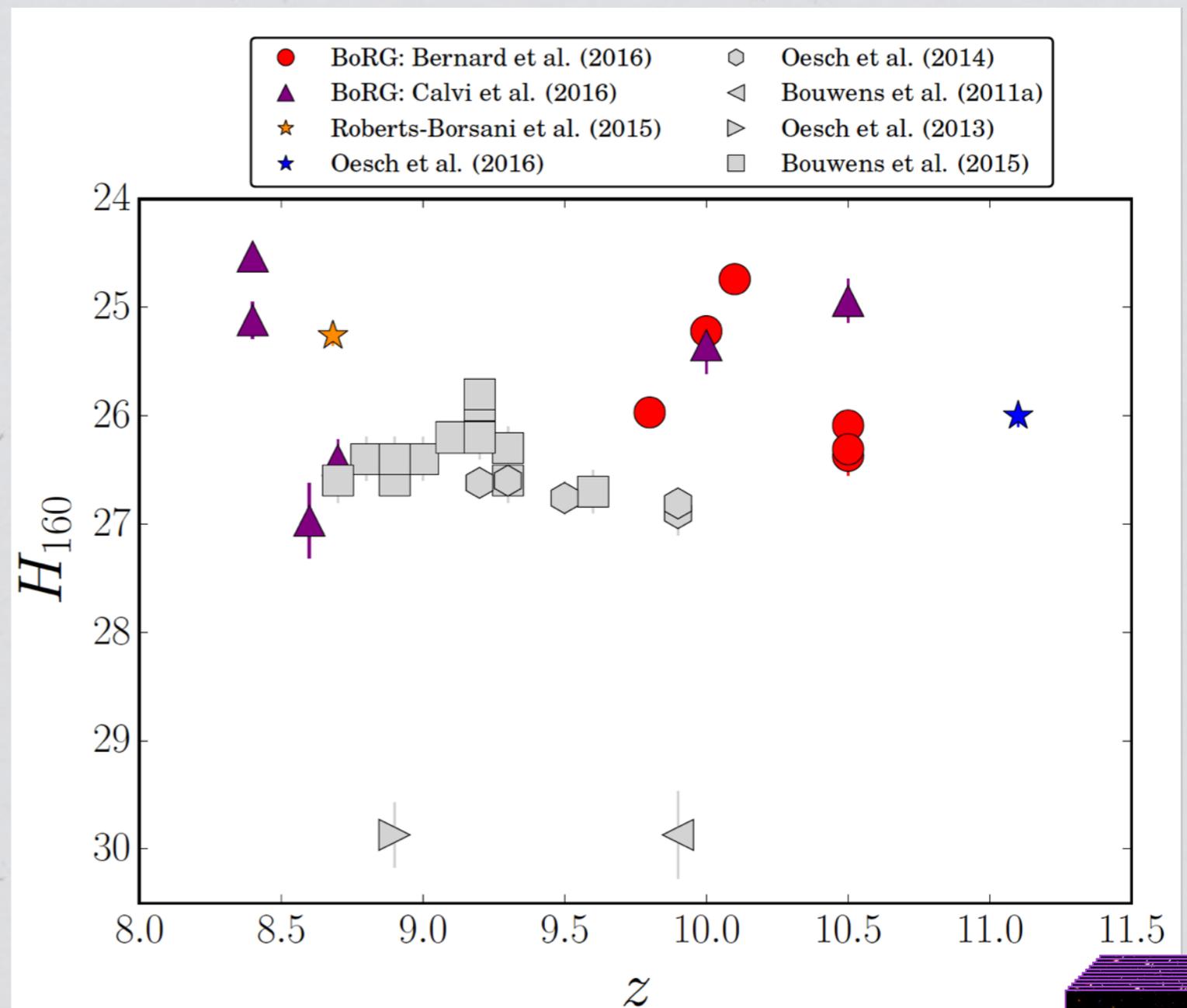


Bright Galaxies at 500 Myr?

★ BoRG $z \sim 10$ candidates are very luminous

★ But comparable to sources with spectroscopic confirmation

★ Keck and Spitzer follow-up approved and underway!

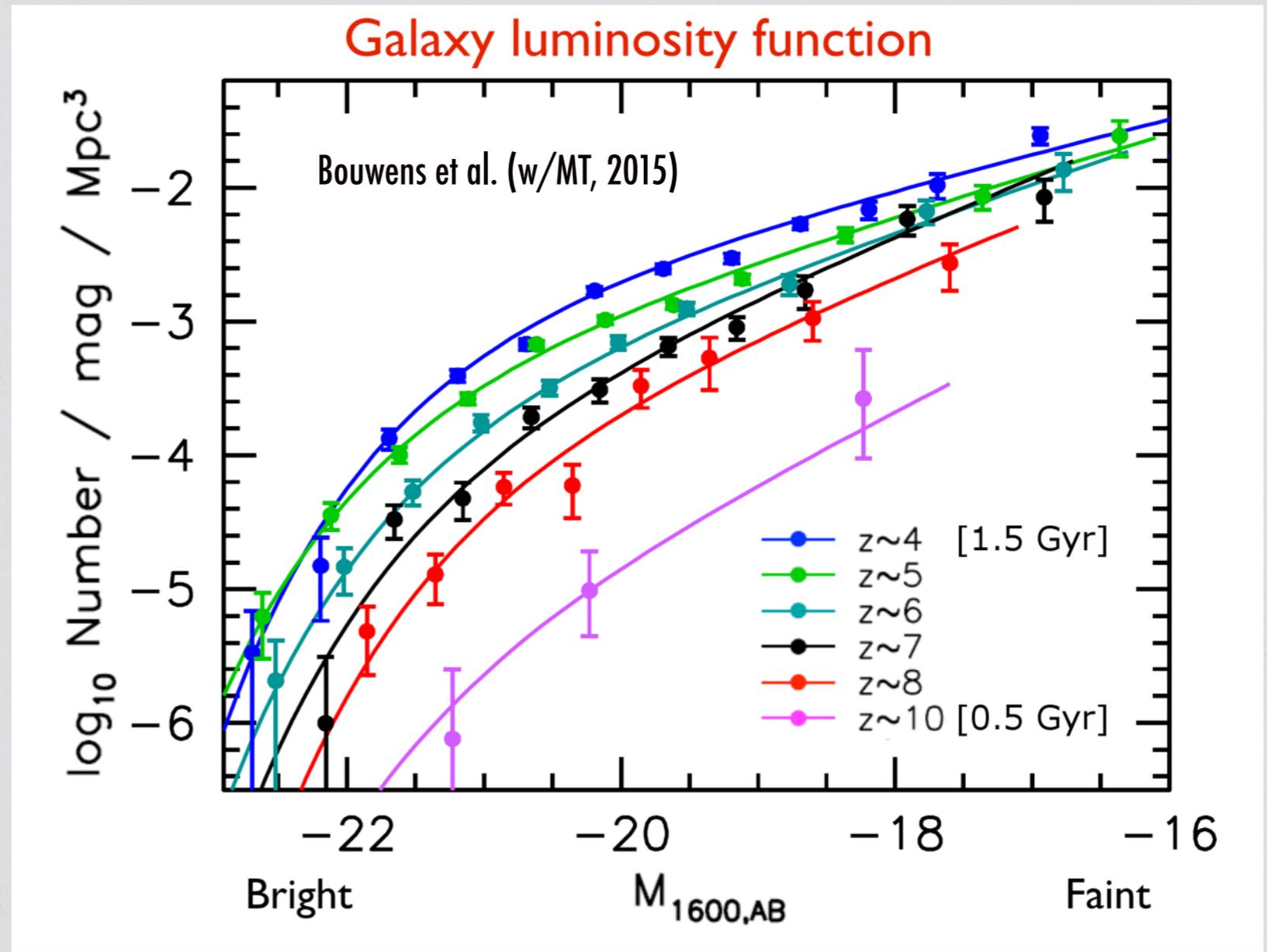


Bernard et al. (submitted)



High-z galaxies: Extending the frontier

★ WFC3/HST:
Galaxies
found at
 $z \sim 7-11$



★ Next step: Characterize their properties

★ Brightest $z > 8$ galaxies: *rare but ideal targets*

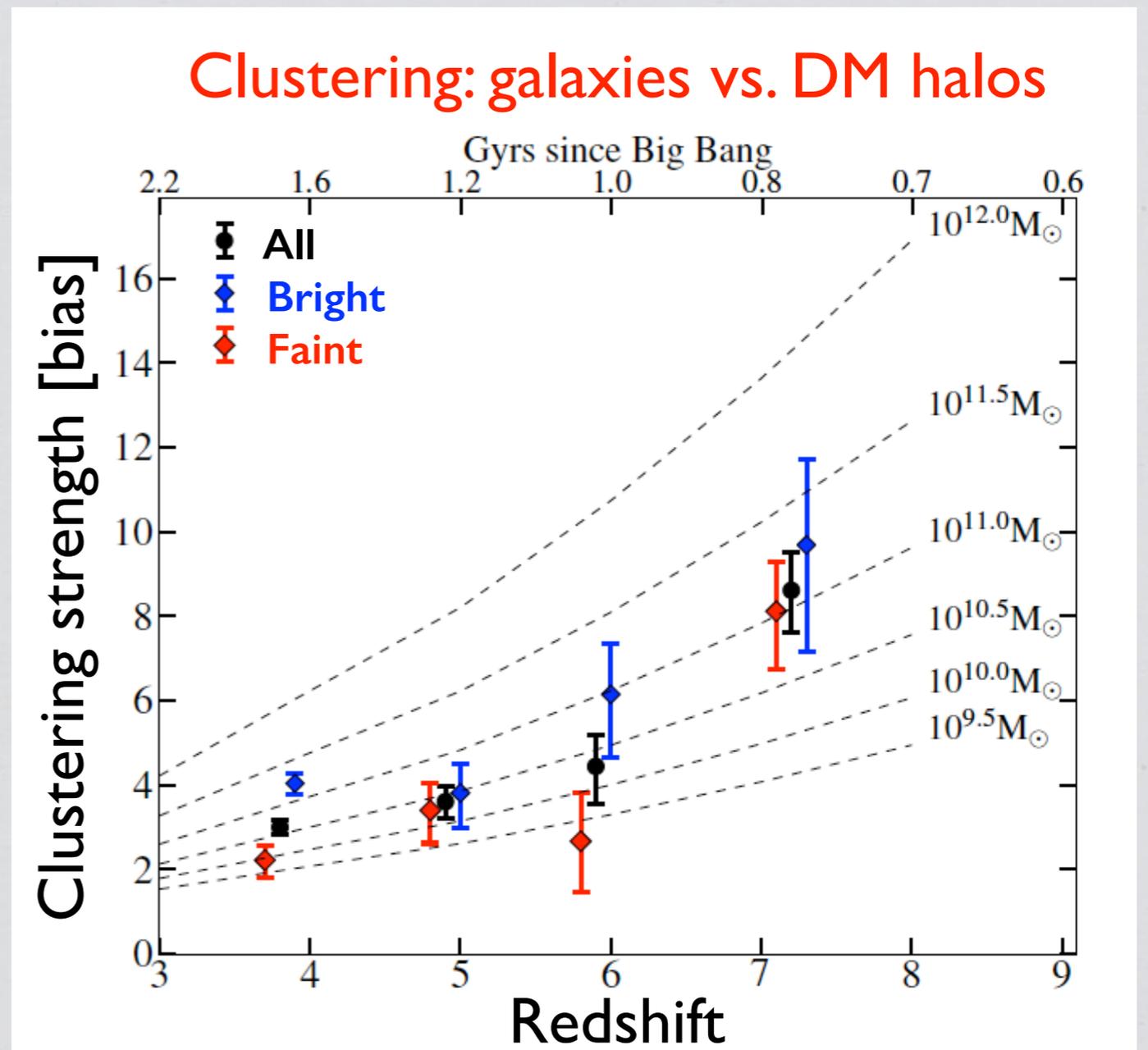
Halo masses from galaxy clustering

Strength of galaxy clustering gives statistical weight of the parent DM halos

★ First measure of galaxy clustering at $z > 7$ (CANDELS data)

★ Observed galaxies have DM halos of $10^{10} - 10^{11} M_{\text{sun}}$

★ Consistent w/theory (e.g. Mason et al. 2015)

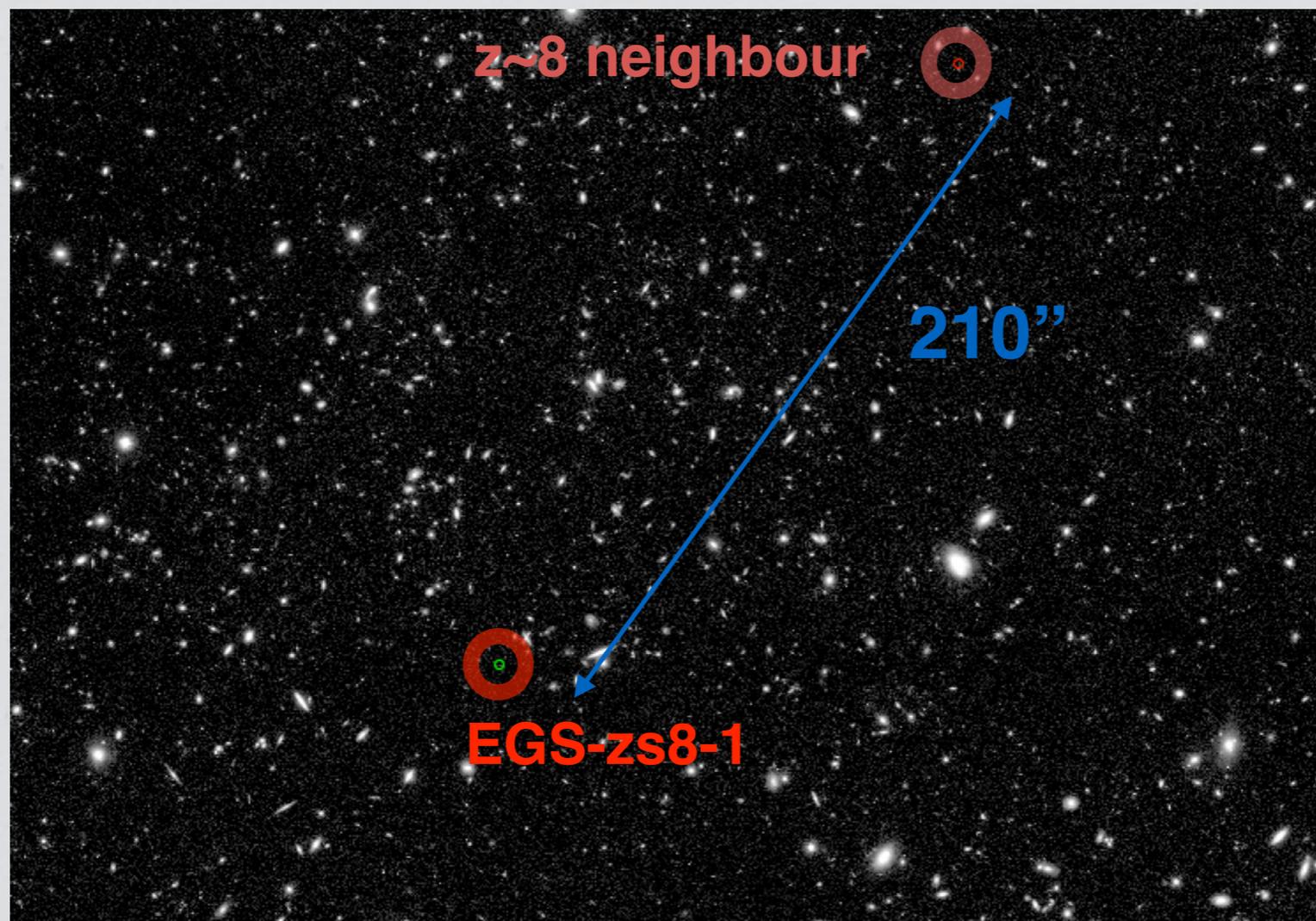


Barone-Nugent, Trenti et al. (2015)

Mysteries of (weaker?) clustering

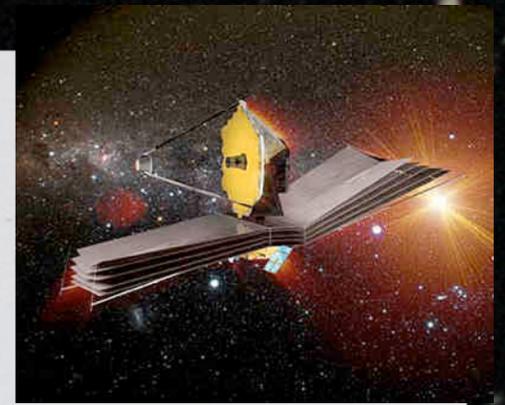
★ Why are the brightest galaxies isolated?

$m_{AB}=25$ galaxy at $z_{spec}=7.7$: No neighbours in $r=200''$ ($m_{AB}<26.3$)



Oesch et al. (2015)

Beyond Hubble: JWST (2018)

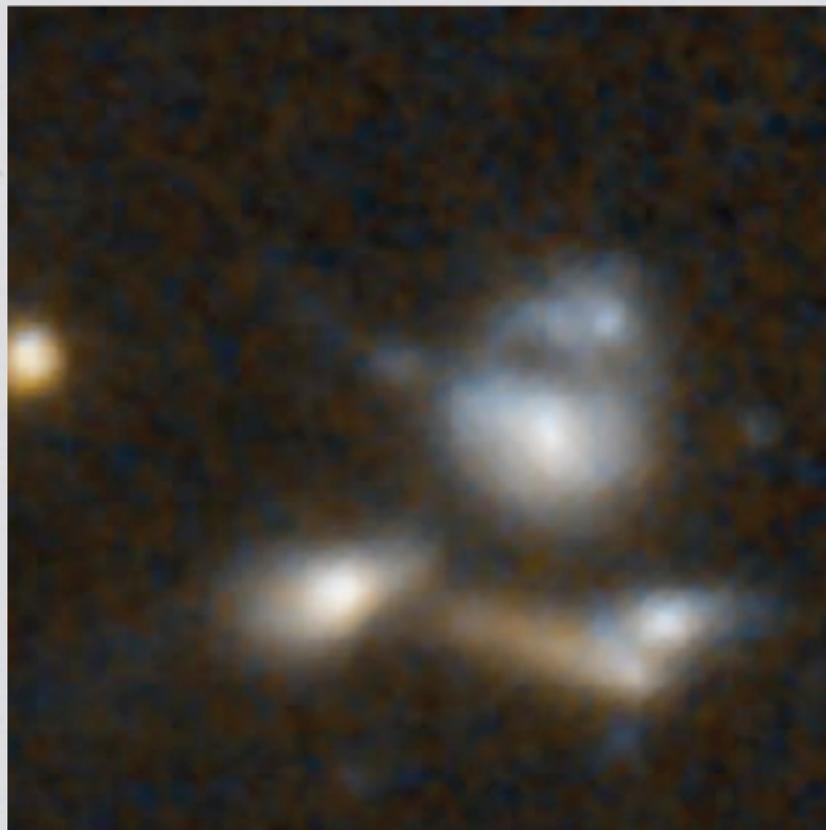


★ James Webb Space Telescope giant leap:

- Gain in sensitivity, resolution, wavelength range [6.5m mirror, $\lambda \sim 0.6\text{-}27 \mu\text{m}$]

Hubble

H band
(1.6 μm)



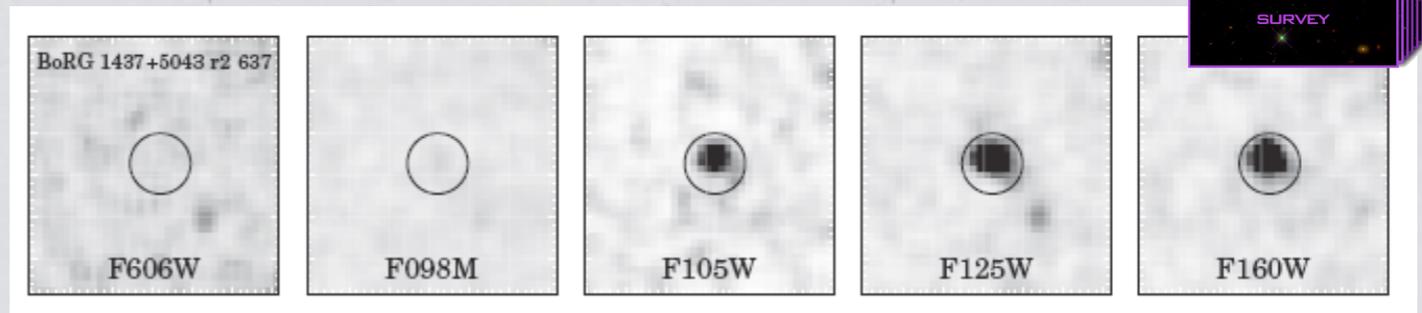
JWST

H band
(1.6 μm)

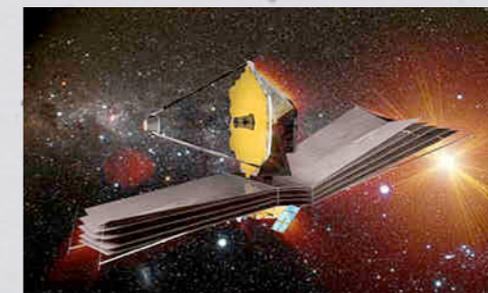


image simulation by
M. Stiavelli, STScI

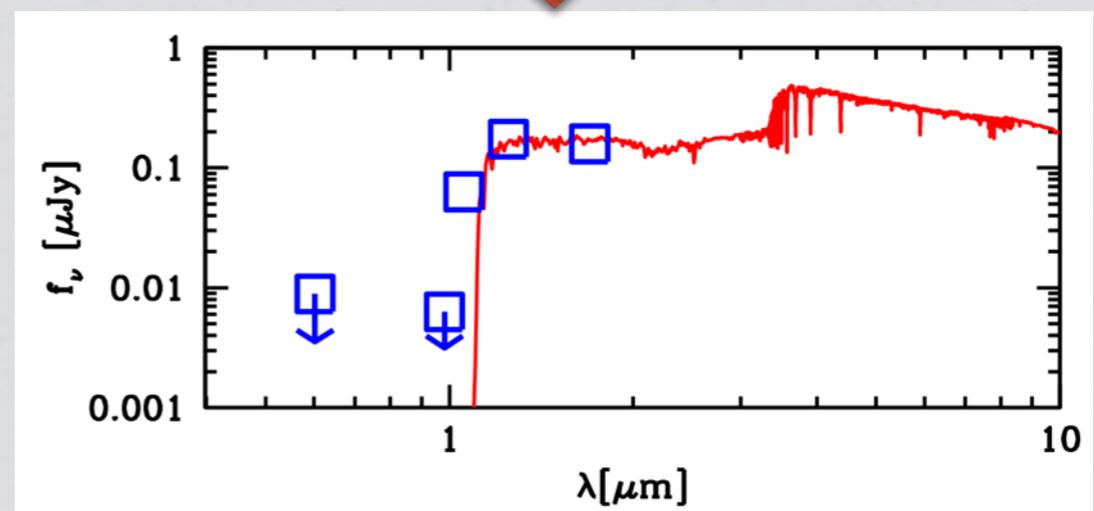
Stellar populations at $z \sim 8-10$



★ ~3h NIRSPEC: Medium-res **continuum spectrum** at restframe optical for $z \sim 8-10$ BoRG galaxies [G395M, R=1000]



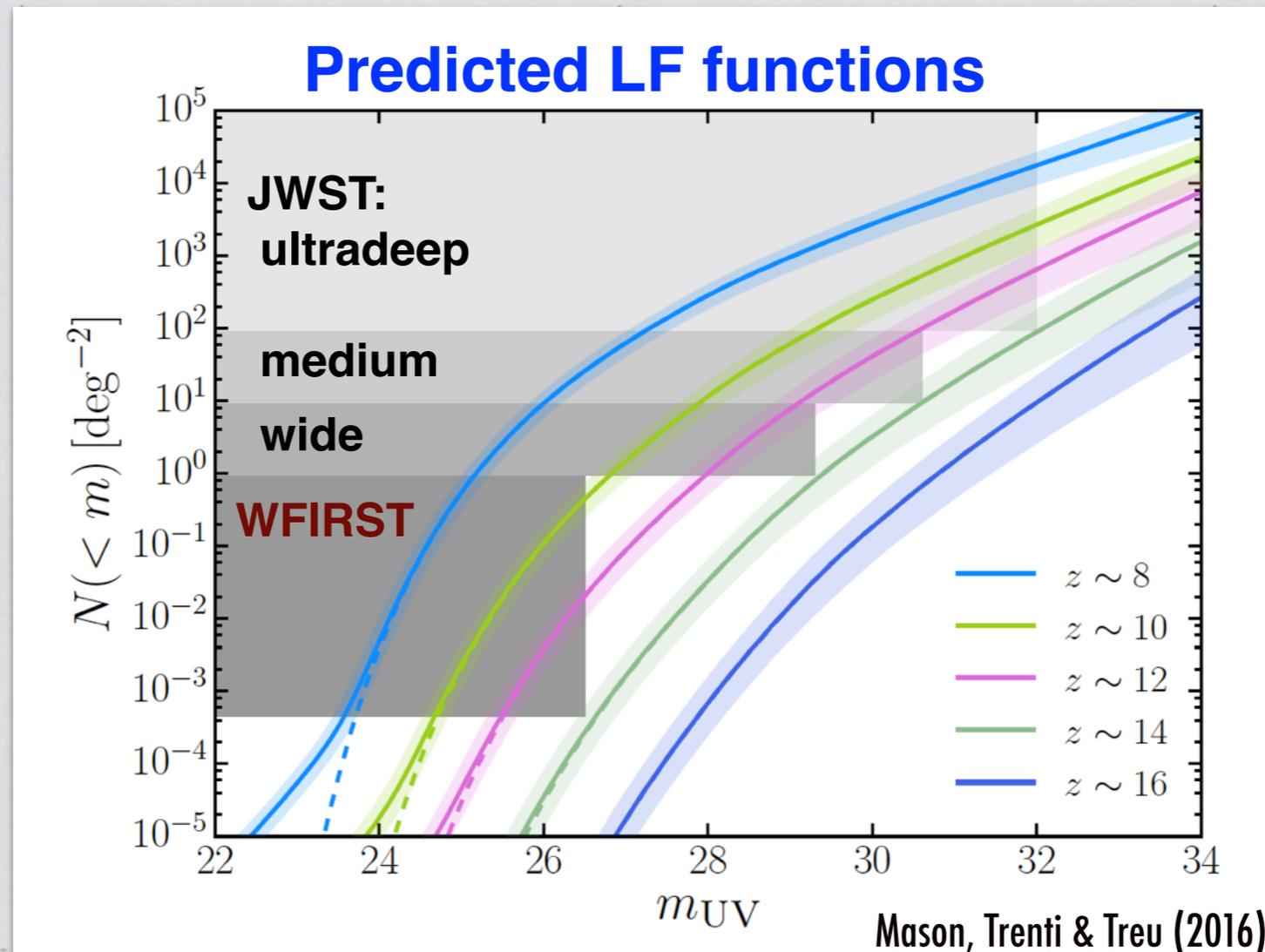
★ Opportunity to identify new targets with HST is now!



Extending the frontier



★ JWST imaging will explore the $z > 10$ Universe

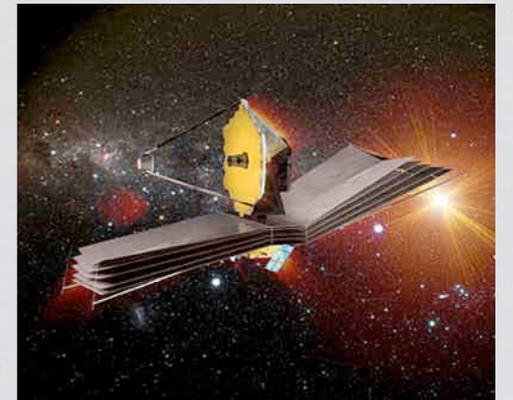


Redshift frontier is reached by wide area surveys:

Opportunity for BoRG-like JWST parallels!



Summary



- Hubble's WFC3 transformed our view of high- z galaxy evolution
- HST parallels have proven potential to explore the $z \sim 10$ frontier
 - Brightest galaxies ideal to advance data-model comparison
- Parallel imaging with JWST will be an amazing opportunity for unprecedented deep observations in the infrared!